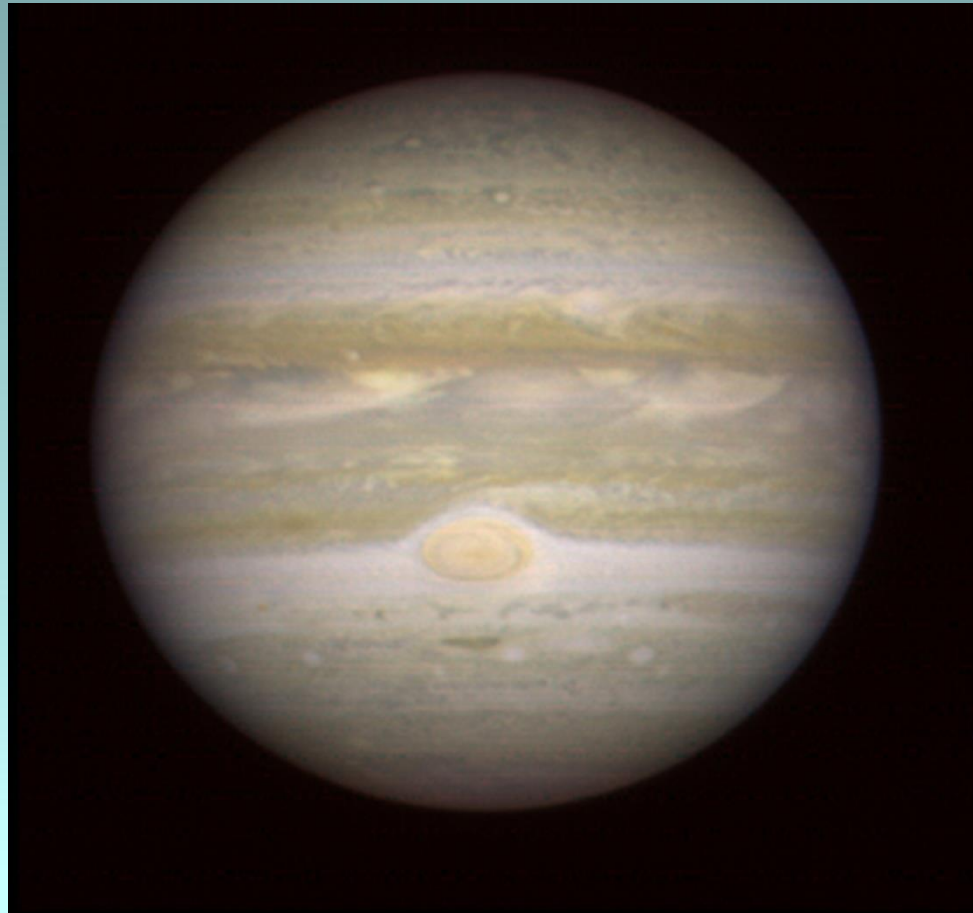


Introduzione alla morfologia cometaria

Giannantonio Milani

Corso teorico pratico sull'osservazione delle comete 7-8 settembre 2019



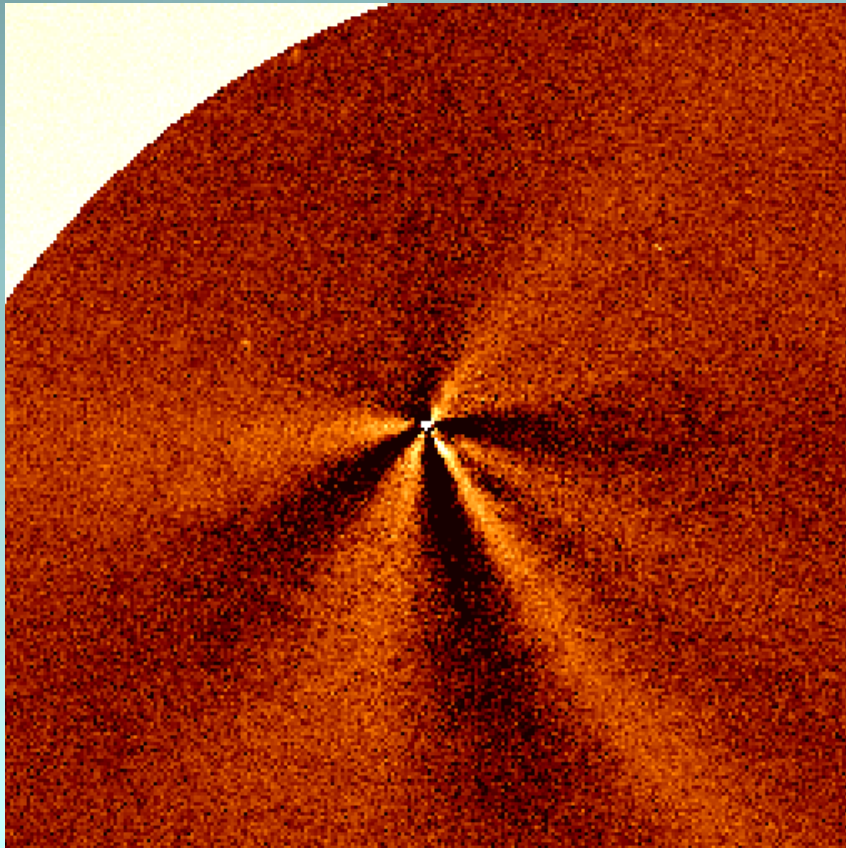


Per molti oggetti del Sistema Solare possiamo fare riferimento a modelli noti che ci consentono di interpretare facilmente le osservazioni



www.jpl.nasa.gov

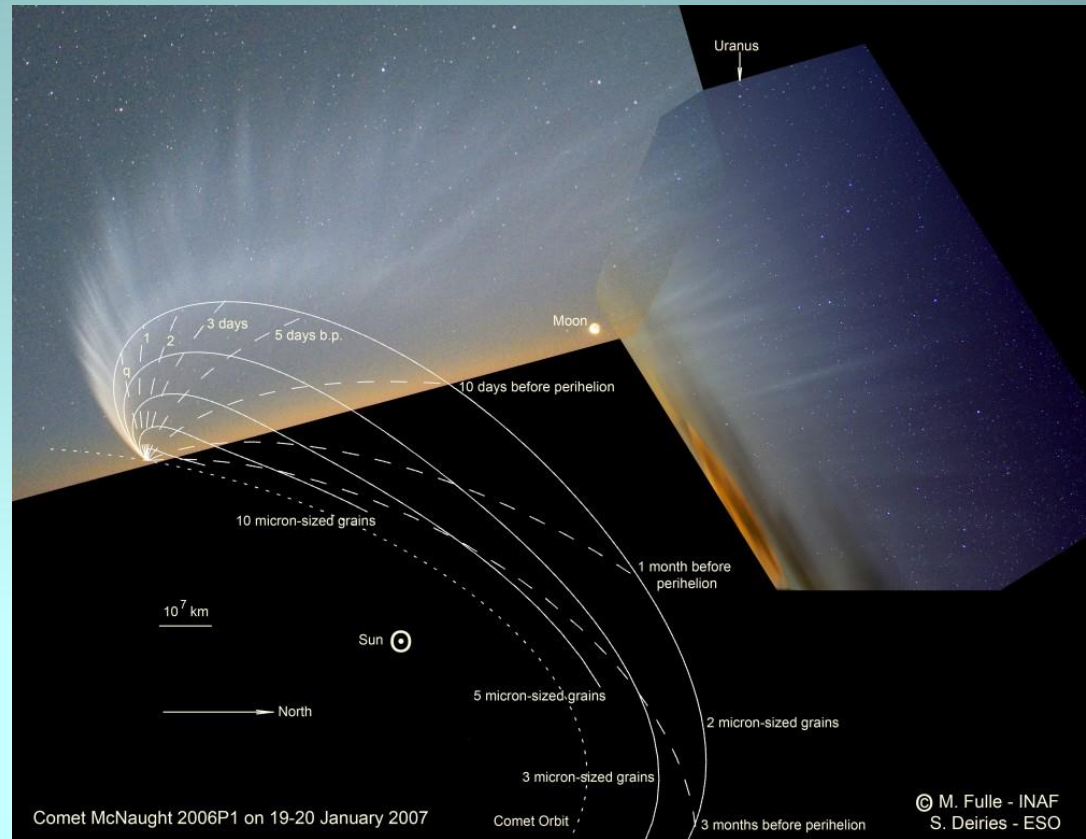
....ma com'è fatta davvero una cometa ?



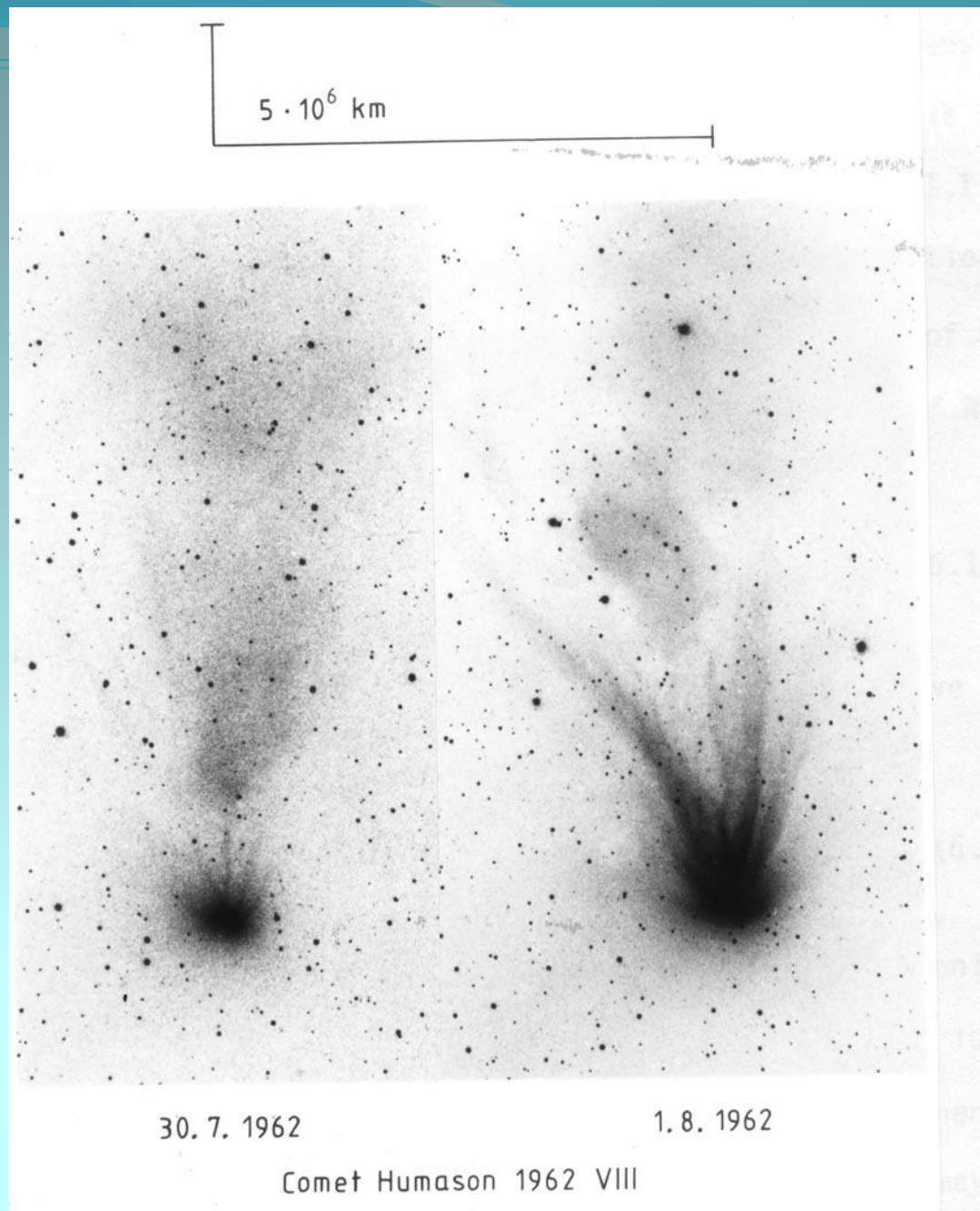
Hale-Bopp: la cometa-porcospino

Le comete offrono un'ampia varietà di forme:

Code a ventaglio e ricurve



aspetti irregolari



comete senza coda

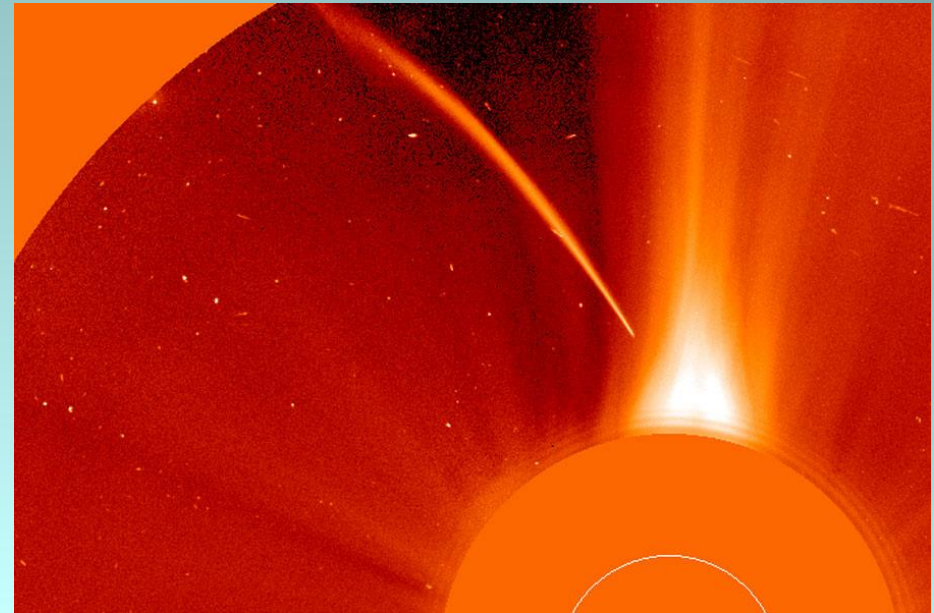
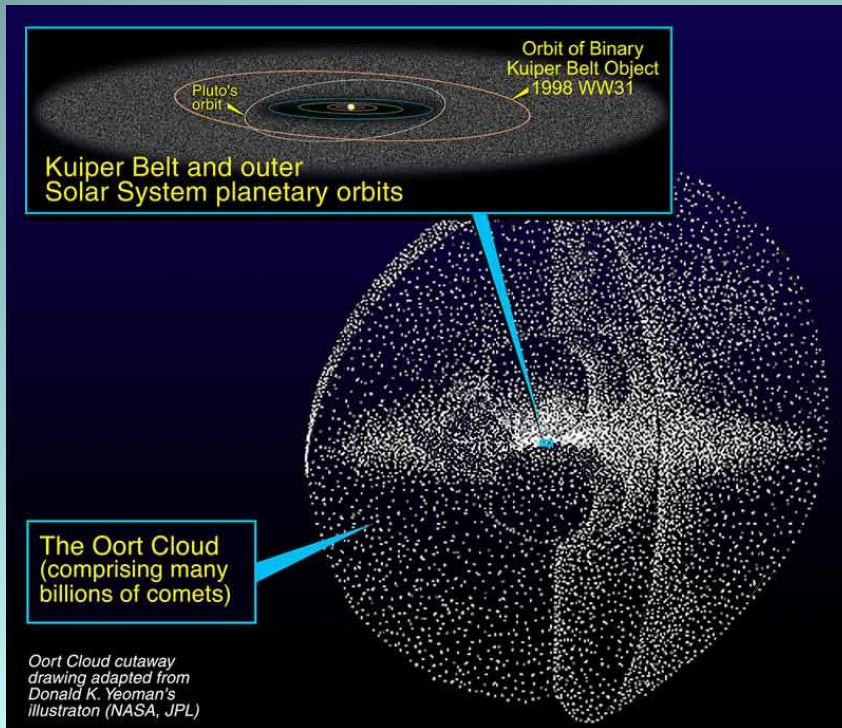


comete ad aculeo (anticoda)



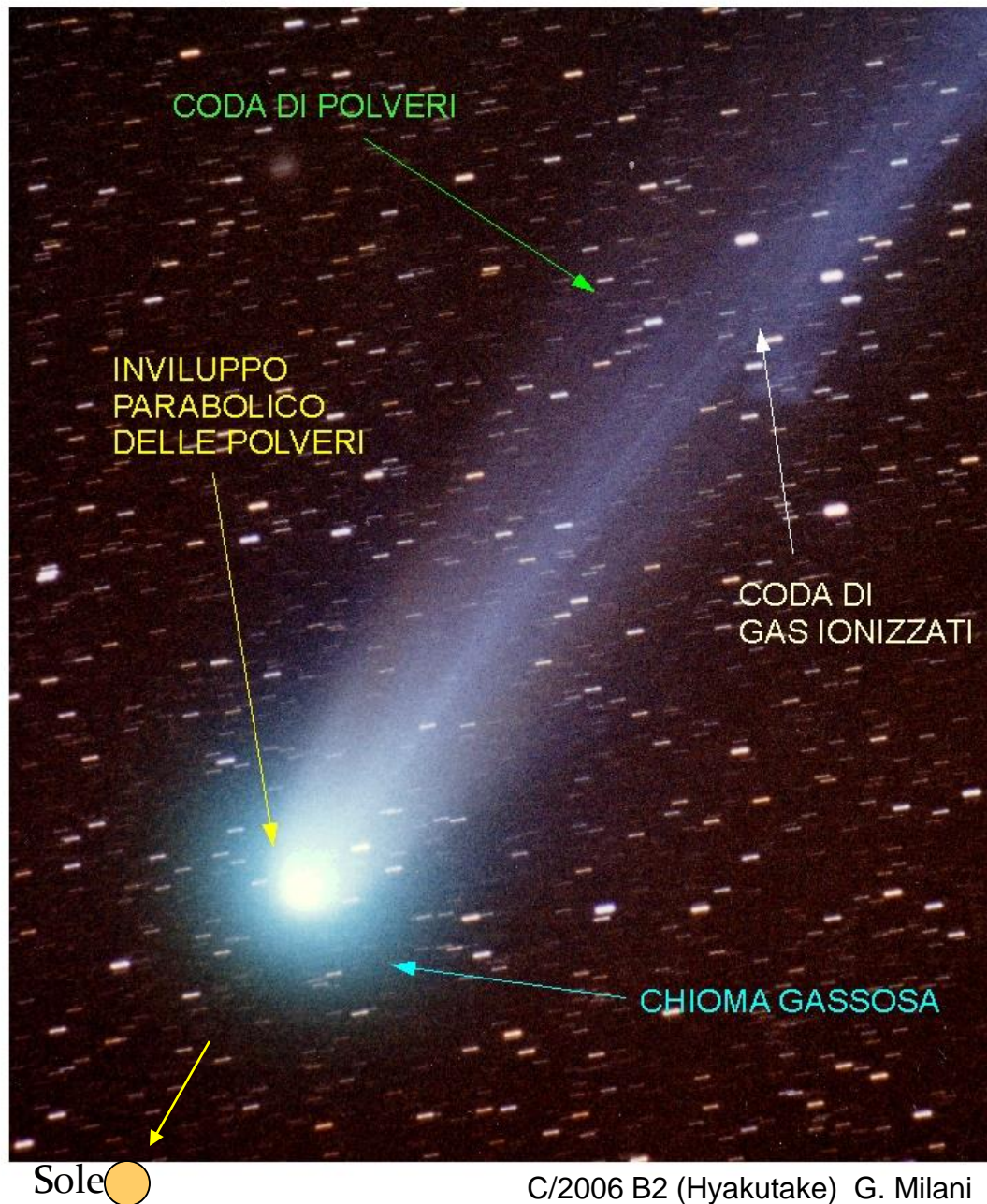
Comet C/2007 N3(LULIN) - 22 febr. 2009 - Erik Bryssinck - Belgium- remote observatory Austr.

Le comete sono gli oggetti che si muovono nel Sistema Solare in un'ampia gamma di distanze: dagli estremi confini fino a sfiorare la superficie del Sole



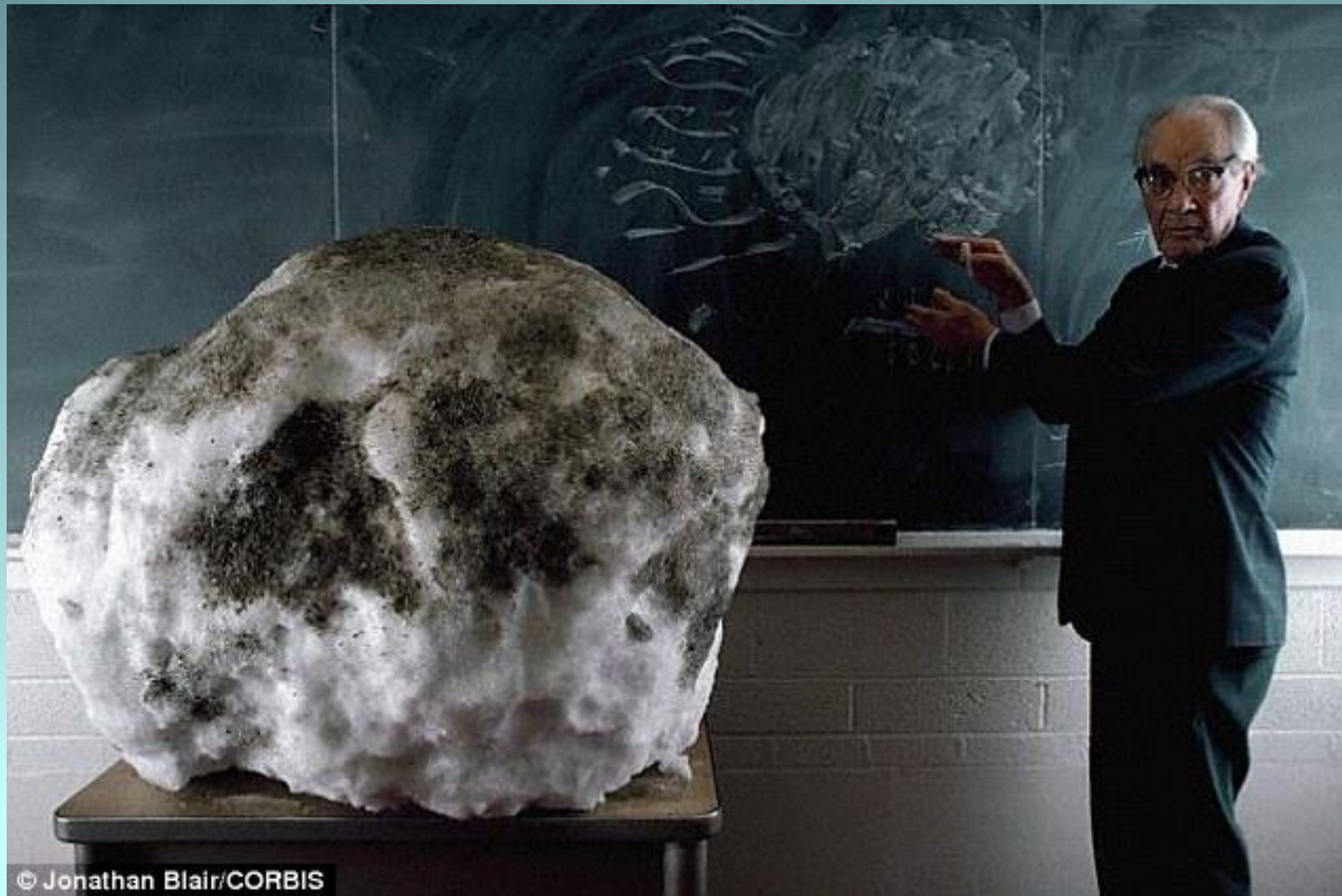
morfologia “classica”

- **chioma gassosa** (nel visibile può estendersi per oltre 200000 km)
- **involuppo di polveri** (esteso circa fino a 50000 km + nella coda)
- **Coda di gas ionizzati**



C/2006 B2 (Hyakutake) G. Milani

DALLA PALLA DI NEVE SPORCA ALLA PALLA DI POLVERI GHIACCiate

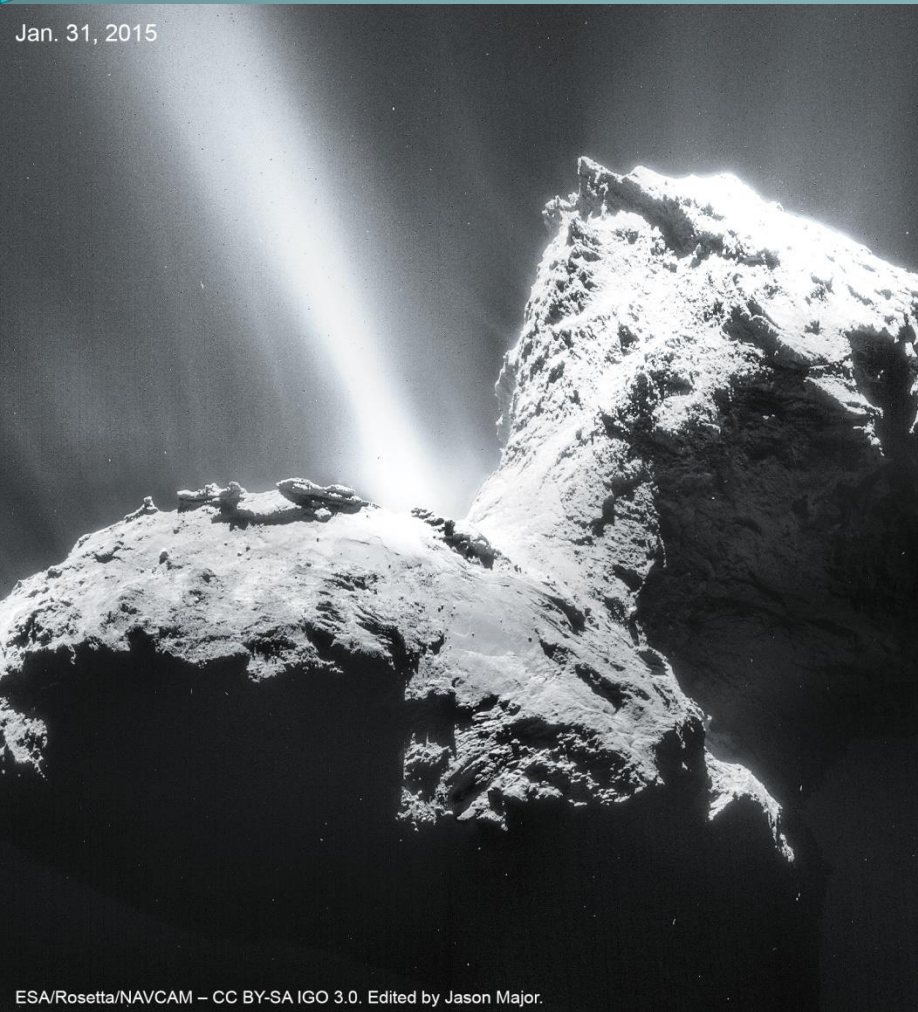


© Jonathan Blair/CORBIS

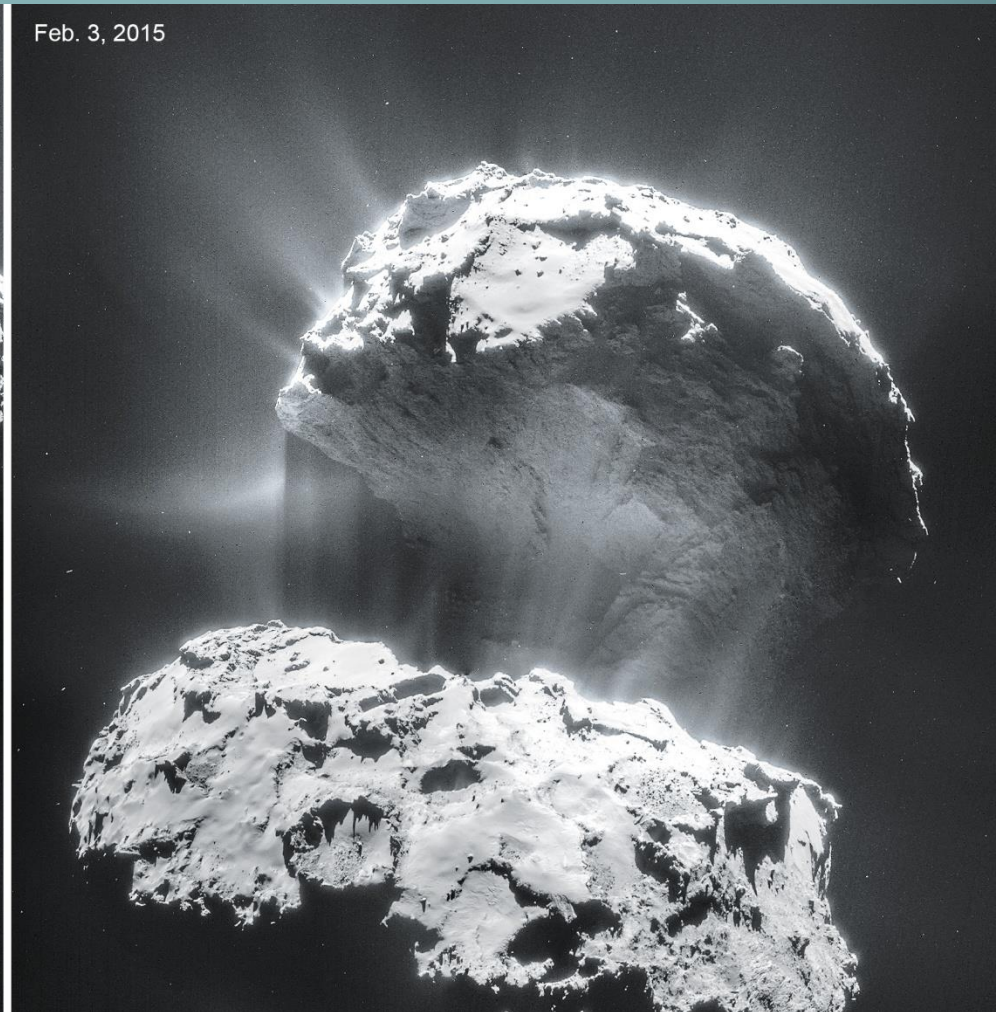
67P - ROSETTA



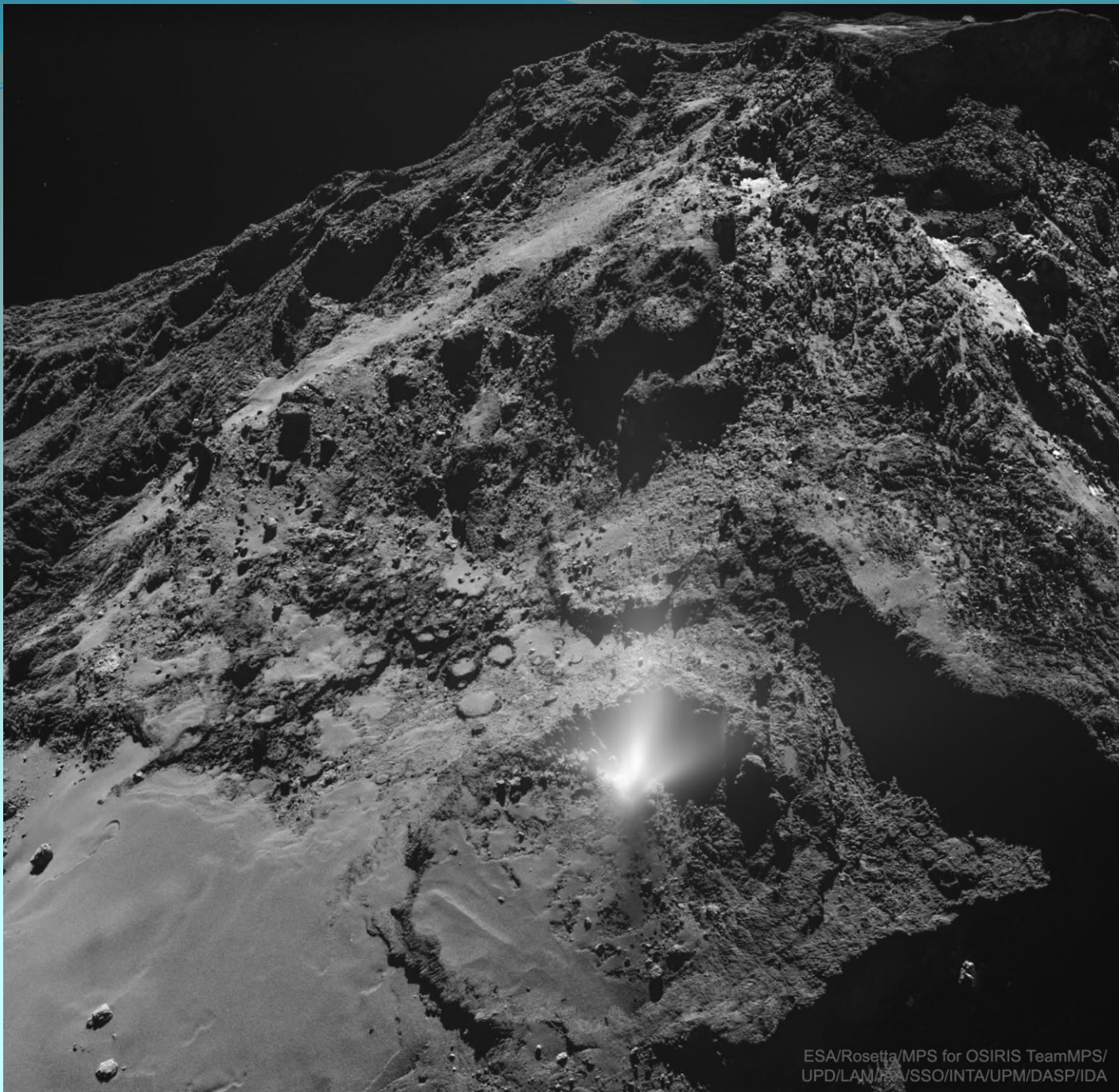
Jan. 31, 2015



Feb. 3, 2015

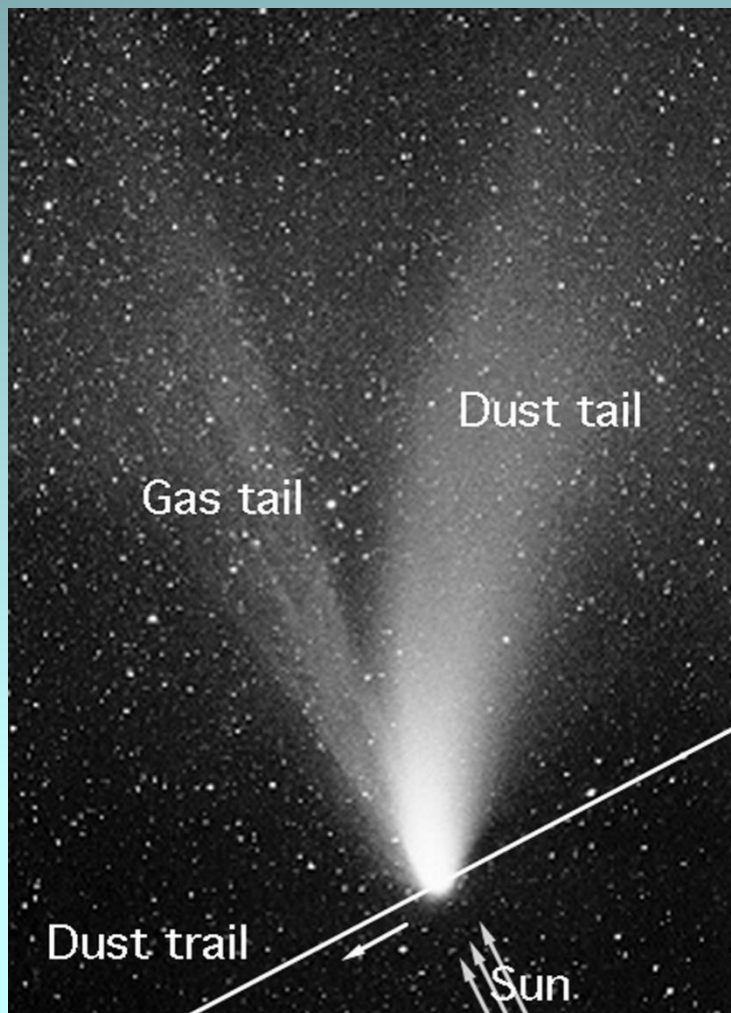


ESA/Rosetta/NAVCAM – CC BY-SA IGO 3.0. Edited by Jason Major.



ESA/Rosetta/MPS for OSIRIS TeamMPS/
UPD/LAM/INAF/SSO/INTA/UPM/DASP/IDA

Non possiamo interpretare una immagine cometaria senza considerare le condizioni geometriche dell'osservazione

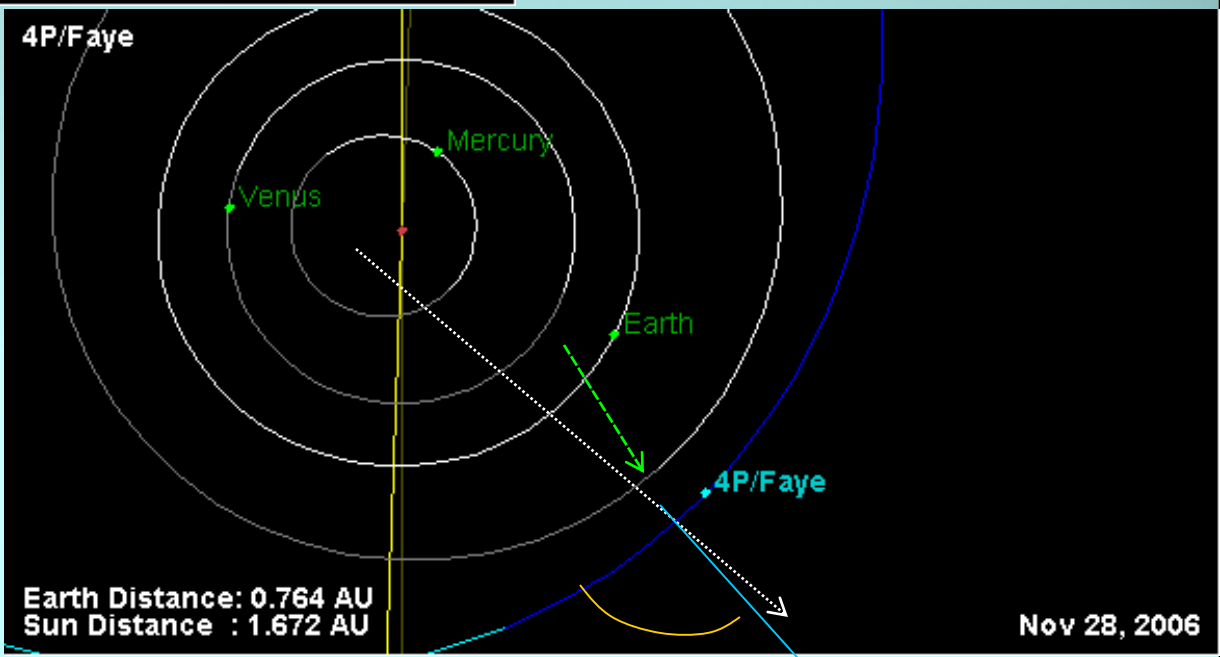
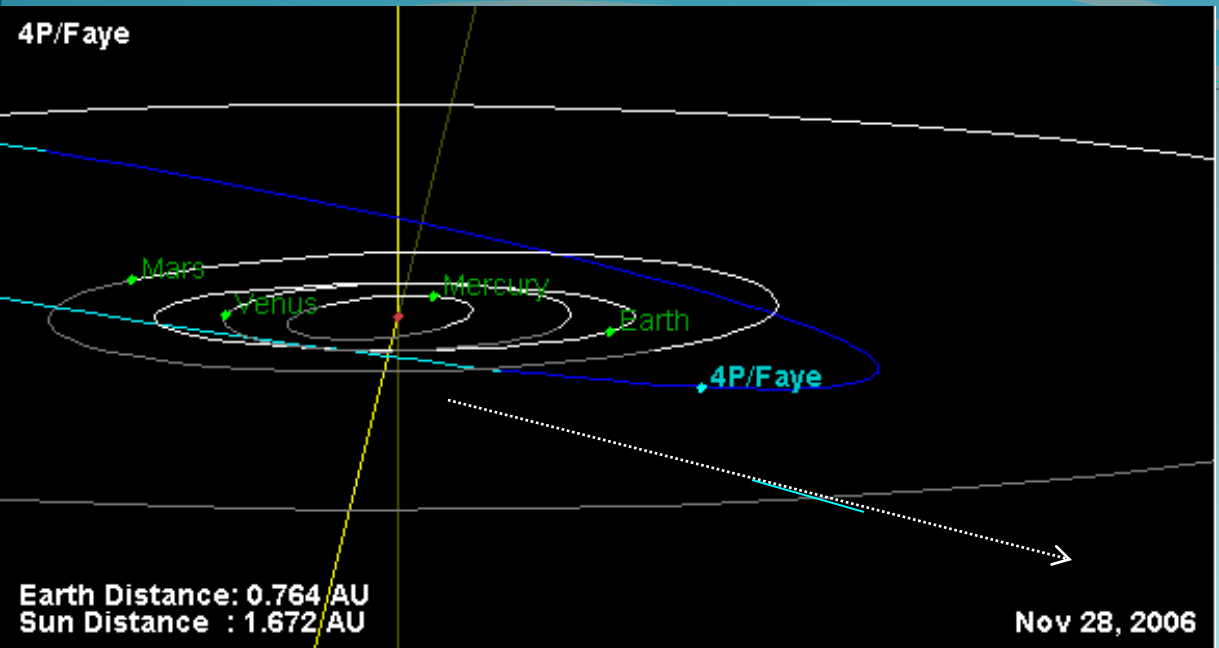


Come è orientata la cometa nello spazio?

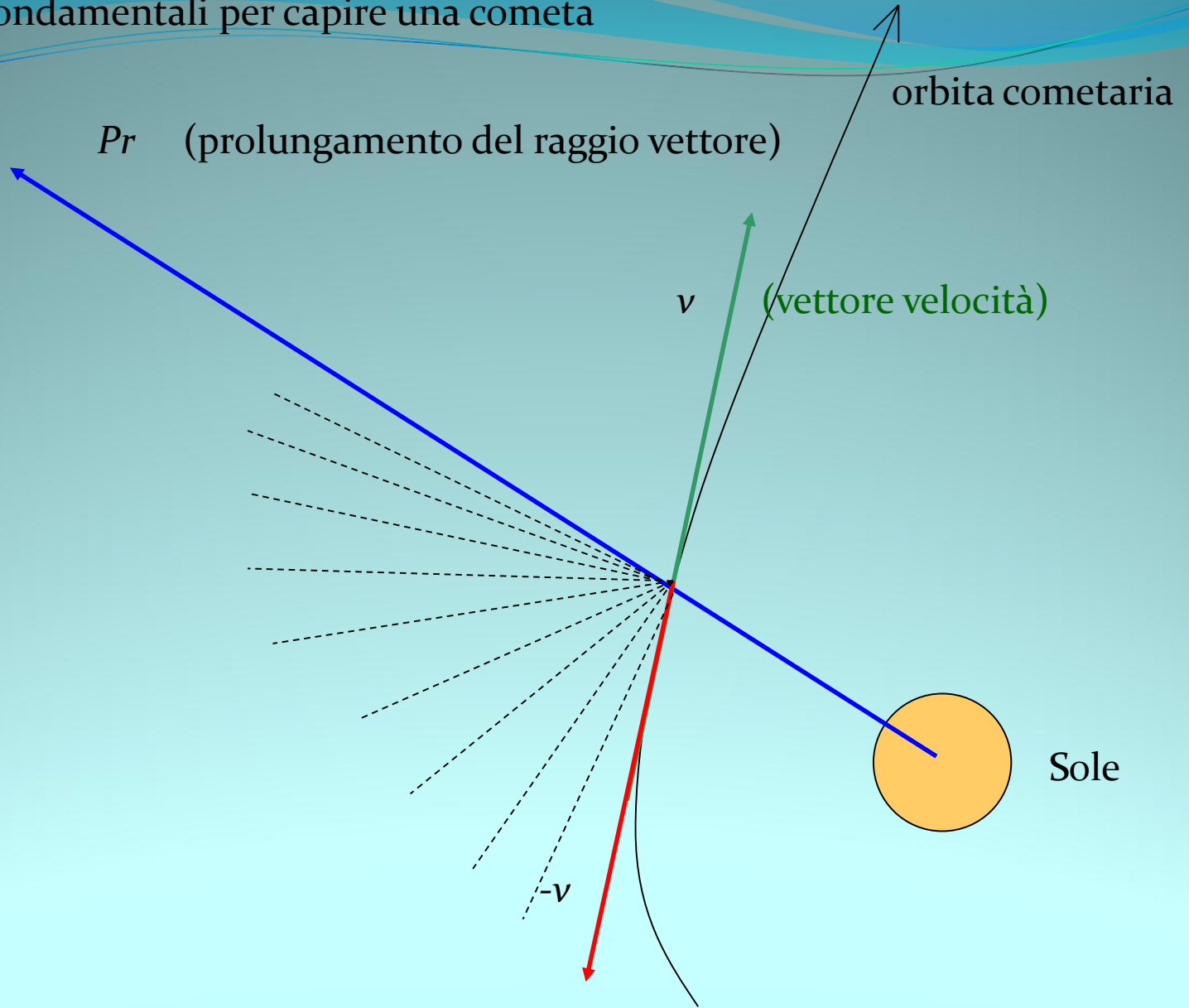
Come la vediamo proiettata sul piano del cielo?

Una prima valutazione può essere fatta mediante l'angolo di fase e la direzione del prolungamento del raggio vettore (congiungente Sole-nucleo)

Come "capire" una cometa?



gli angoli fondamentali per capire una cometa



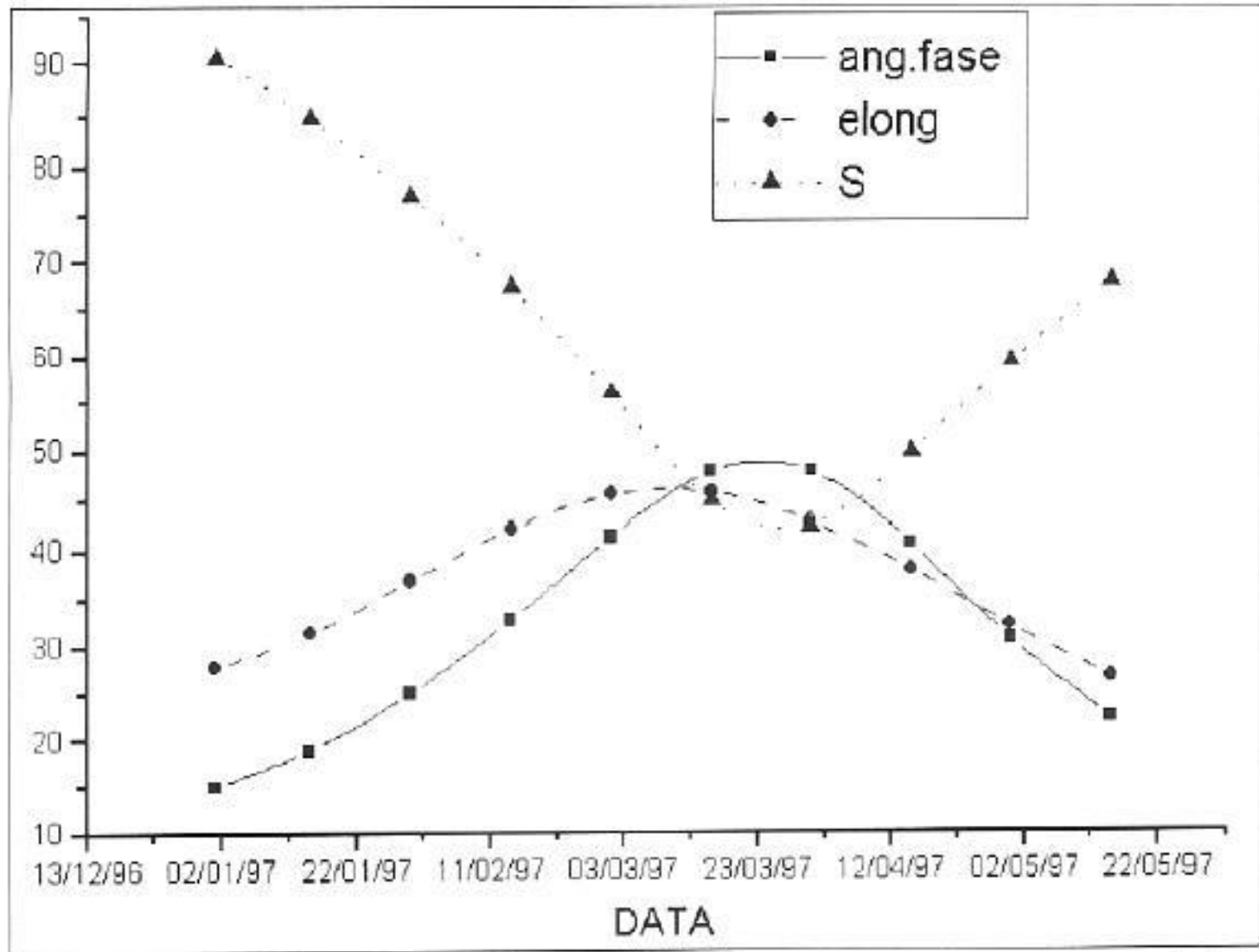
...alcuni degli angoli fondamentali

- P_s - prolungamento del raggio vettore
- angolo di fase - Terra-Sole-Cometa
- S - angolo tra piano del cielo e piano orbitale della cometa

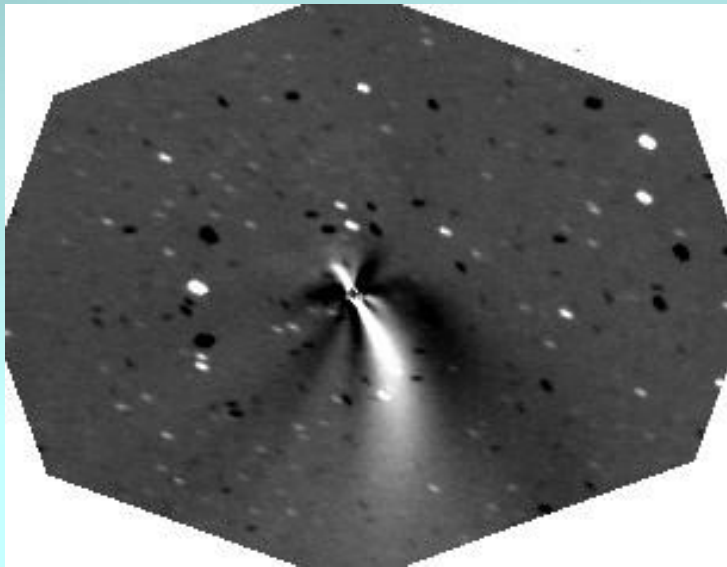
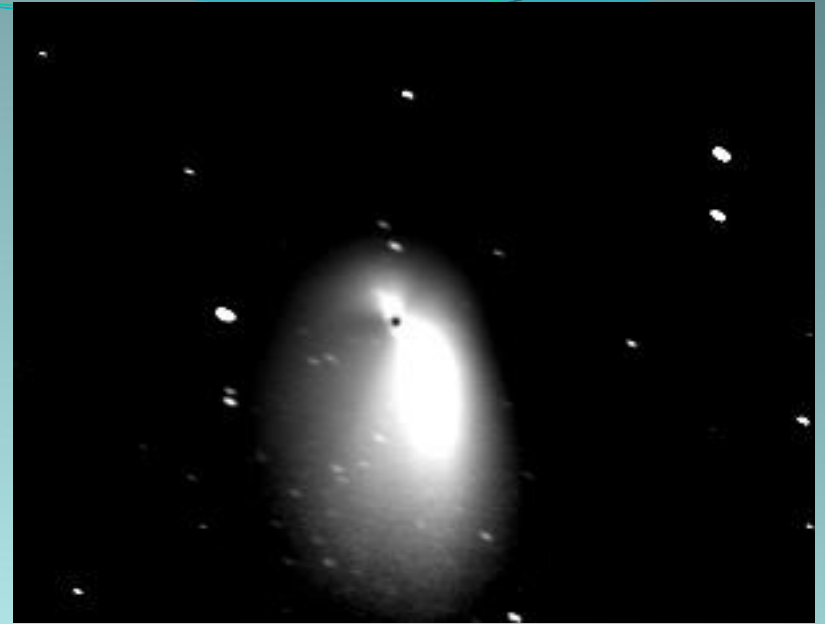
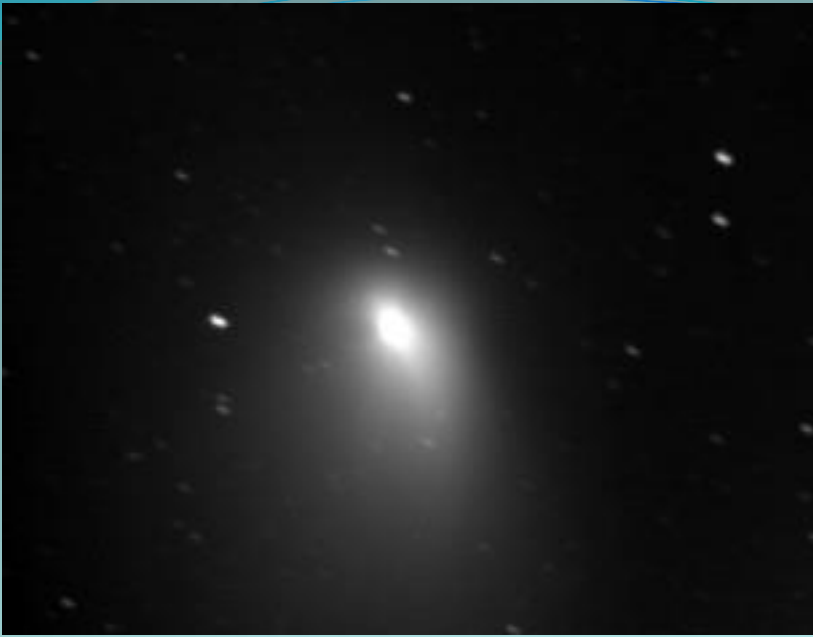
$$\tan(p_s) = \frac{-\sin(\alpha - \alpha')}{\cos \delta \operatorname{tg} \delta' - \sin \delta \cos(\alpha - \alpha')}$$

$$\cos(S) = -\cos i' \sin \delta + \sin i' \cos \delta \sin(\alpha - \Omega')$$

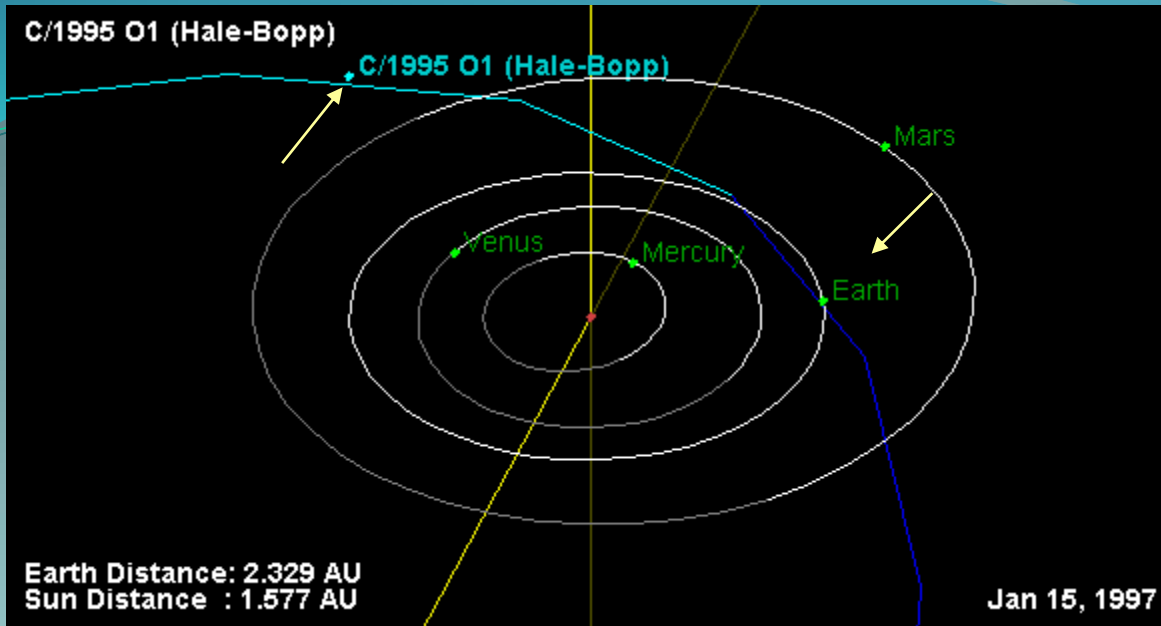
Cometa Hale-Bopp



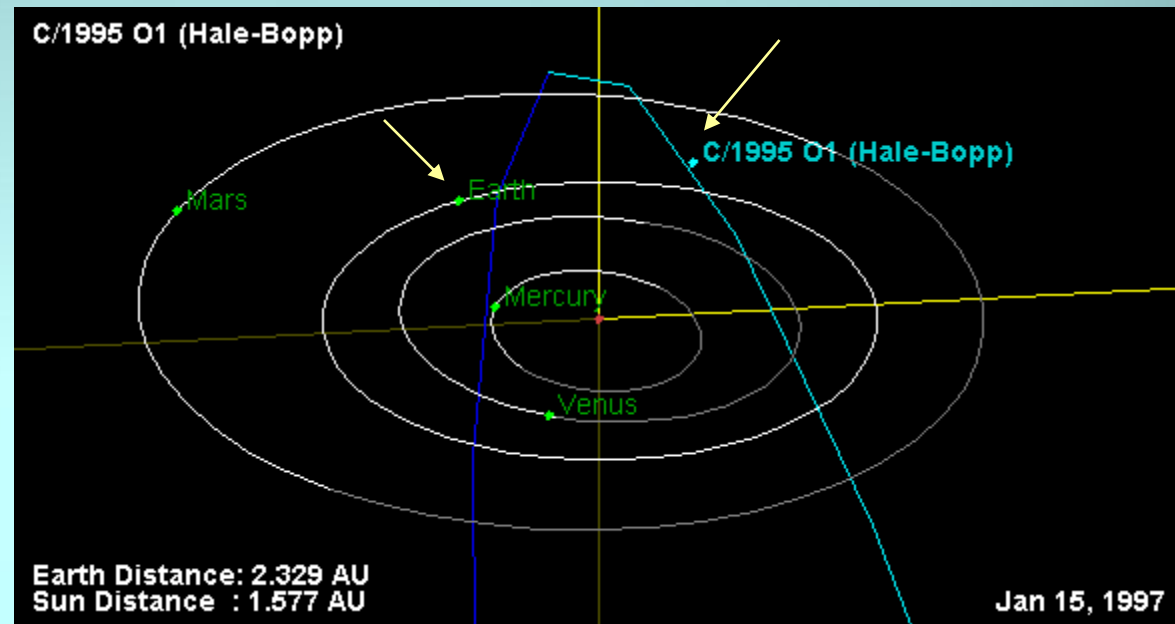
....analizziamo un caso reale



Cometa Hale-Bopp 17 gennaio 1997

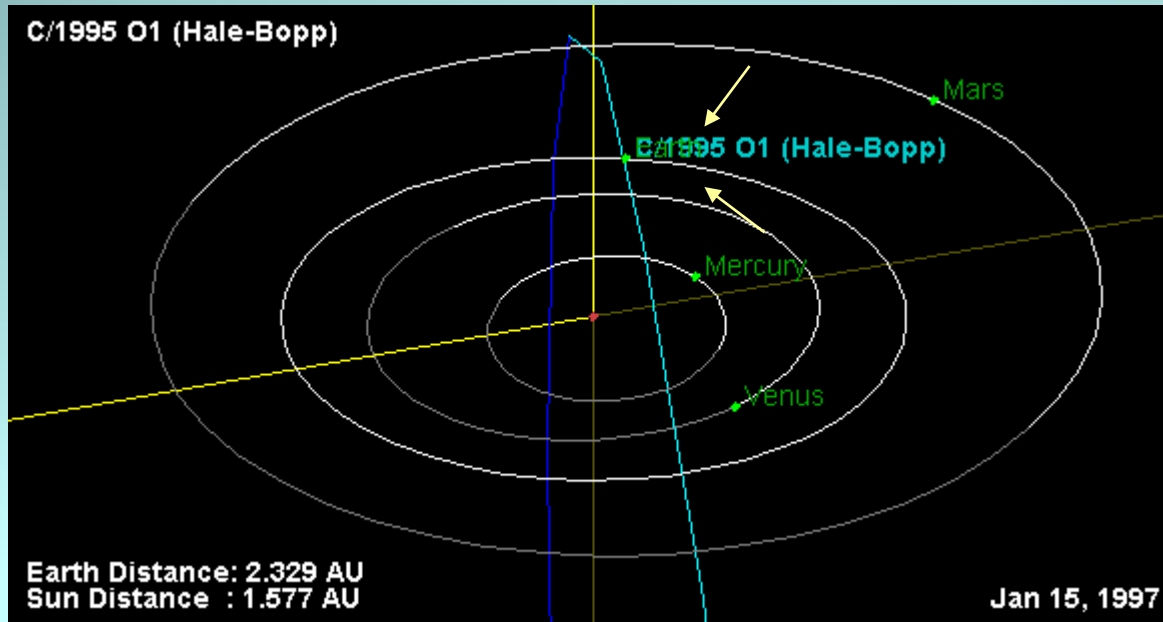
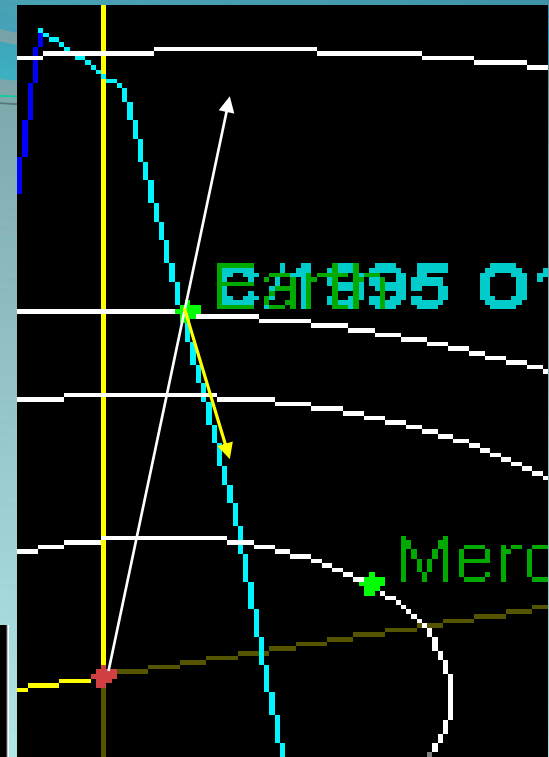
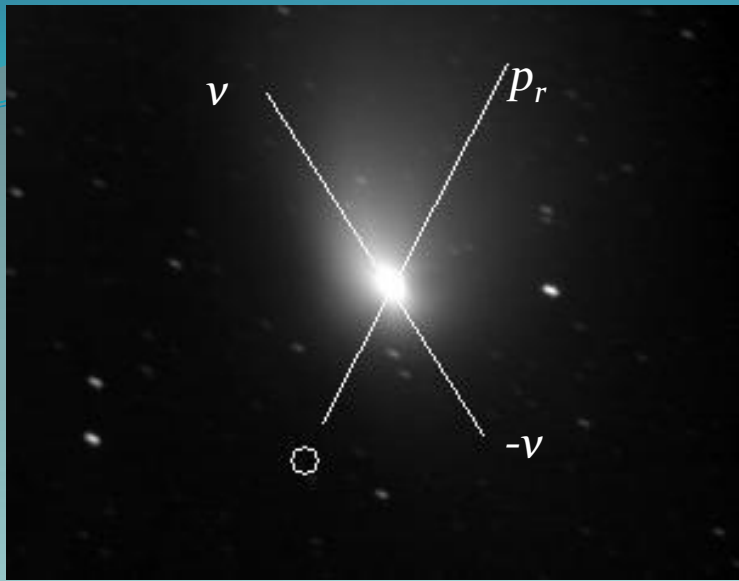


un metodo approssimativo



Per una valutazione “spiccola”:
Orbit diagrams

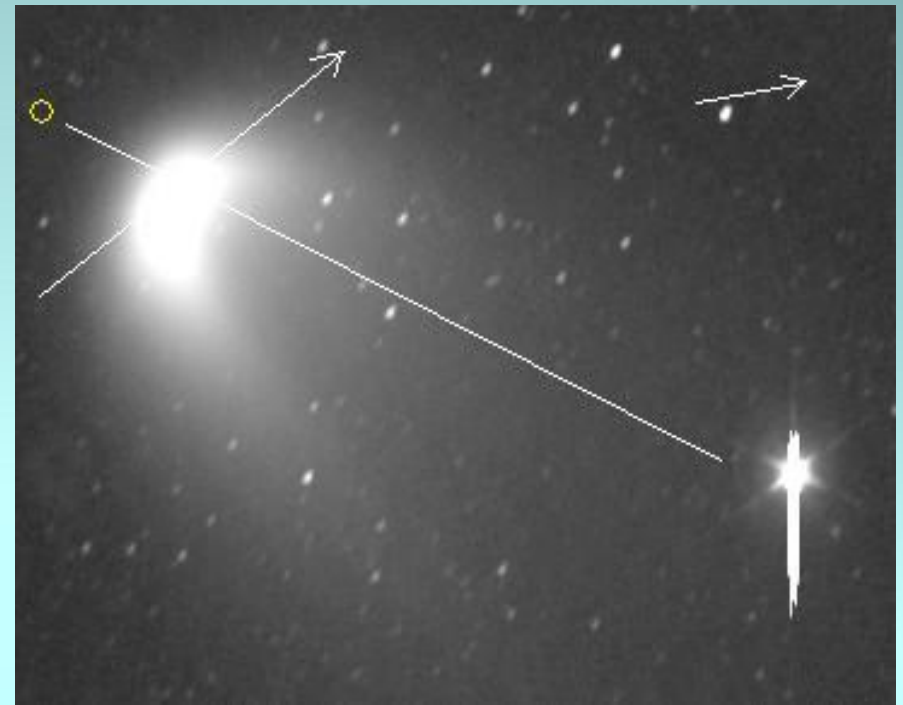
http:
[//neo.jpl.nasa.gov/orbits/](http://neo.jpl.nasa.gov/orbits/)



n.b. la rappresentazione è rispetto al piano dell'eclittica



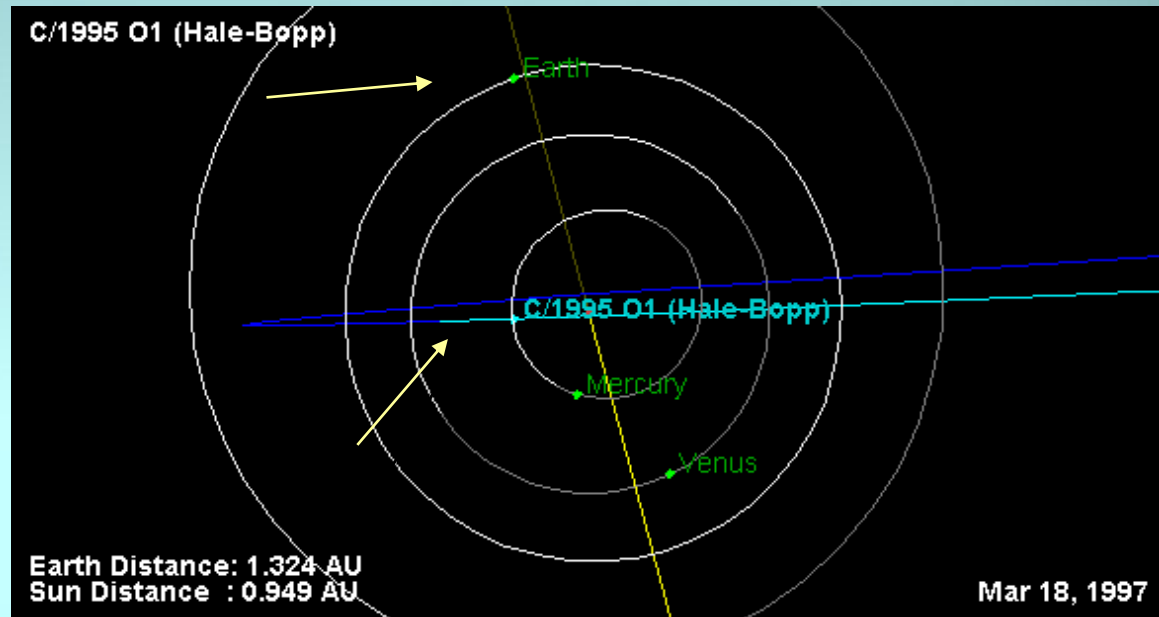
Hale-Bopp 8 febbraio





Hale-Bopp fine marzo:
Terra all'incirca perpendicolare al piano
dell'orbita

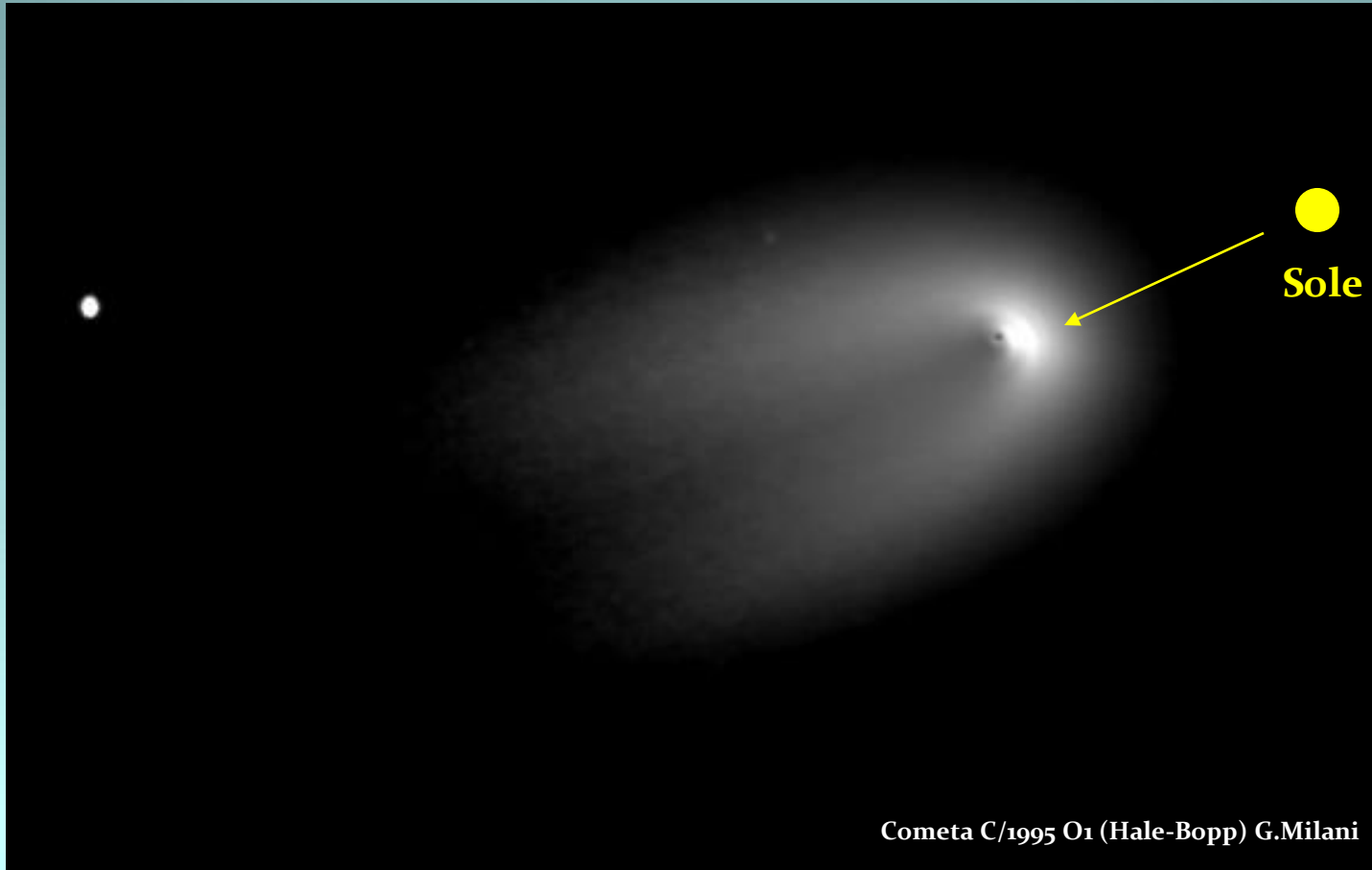
In questo periodo abbiamo osservato al meglio
la coda, le sue strie, ed anche gli aloni
concentrici nella chioma





G.V.Schiaparelli Observatory VARESE Italy
C14 Schmidt Camera & Kodak Film TP 6415 (8 min exposure)
2005, 07 January - 21.19 U.T. - Giorgio Dalla Via

L'involuppo parabolico delle polveri



modello a "fontana" (Eddington), descrive il moto di grani di polvere (moto uniformemente accelerato)

la posizione di un grano di polvere in un sistema di assi x,y,z , è data da

$$x = V_d \sin(\theta) \cos(\phi) t$$

$$y = V_d \sin(\theta) \sin(\phi) t$$

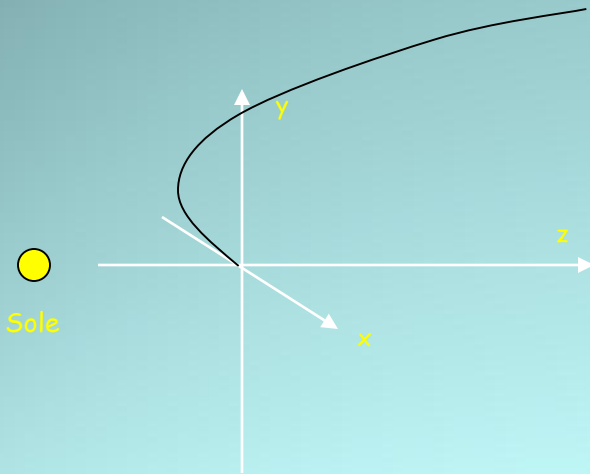
$$z = V_d \cos(\theta) t + 1/2 g_r t^2$$

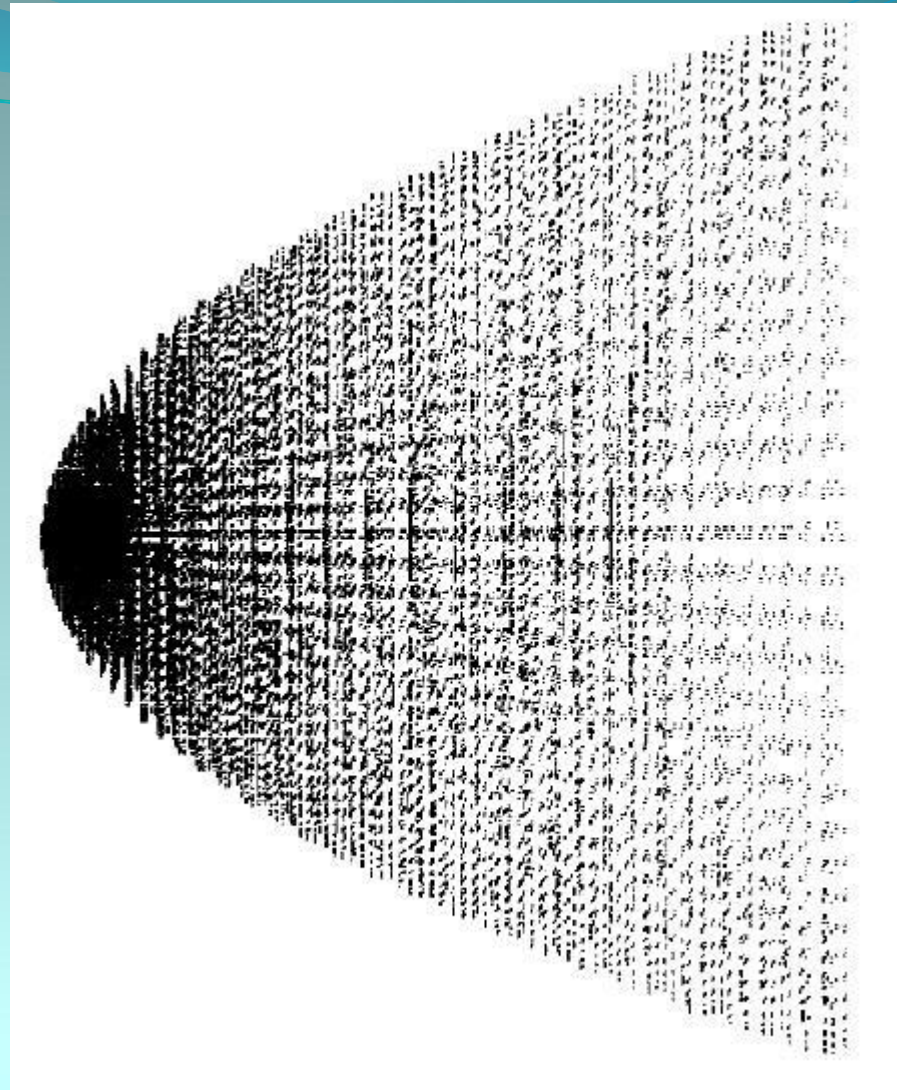
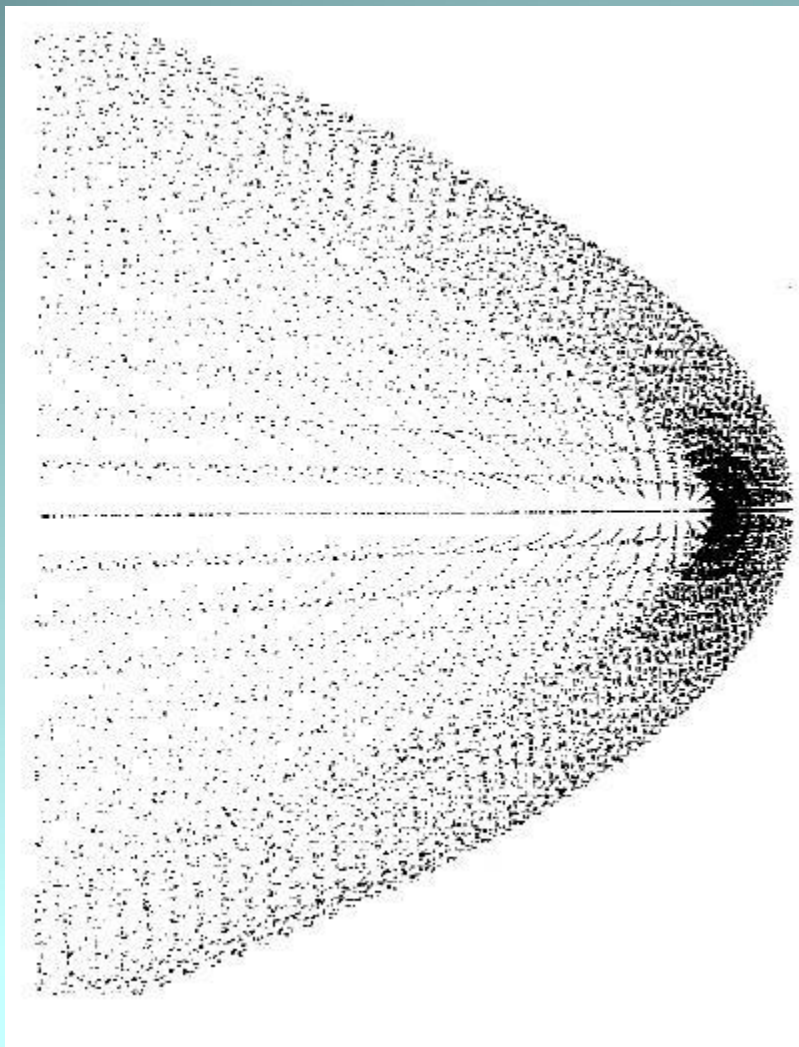
V_d = velocità di emissione del grano

θ, ϕ = angolo di emissione del grano

g_r = accelerazione (repulsiva)

t = tempo





Simulazioni della cometa Hale-Bopp

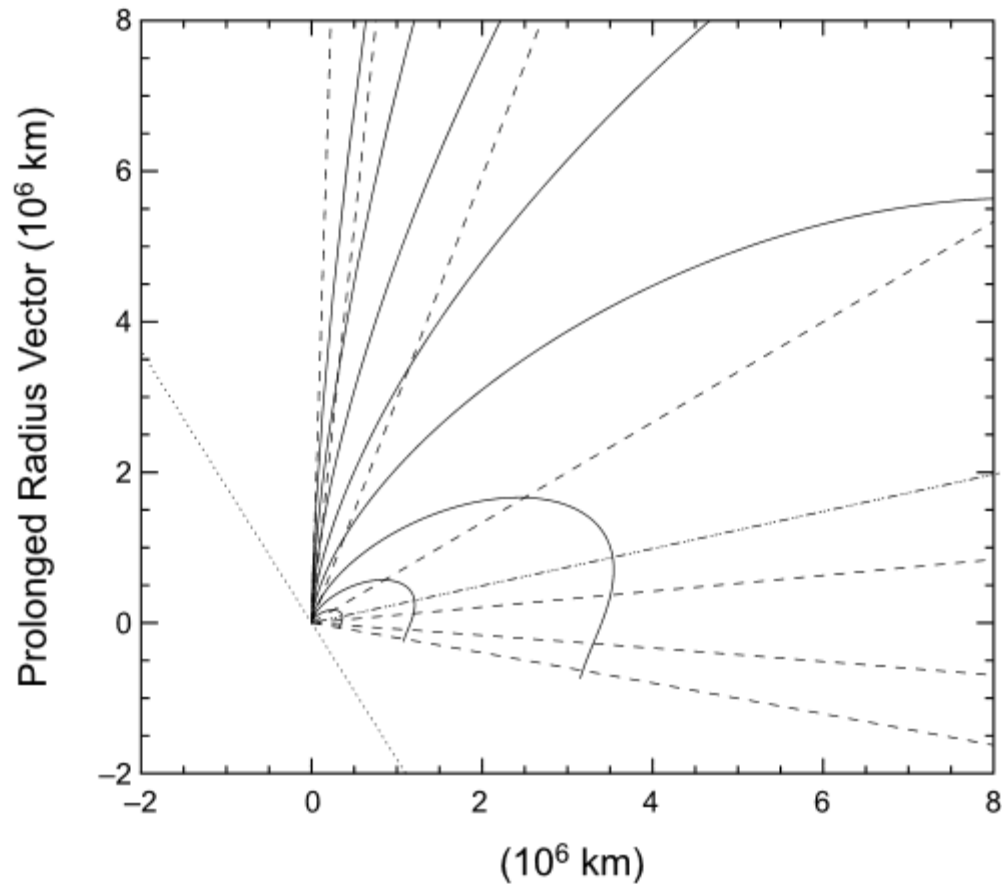


Fig. 1. Synchro-syndyne network for Comet Hale-Bopp 1995O1 on 3 January 1998, when a neck-line was observed during the Earth crossing of the comet orbital plane (Fig. 2). The Sun was

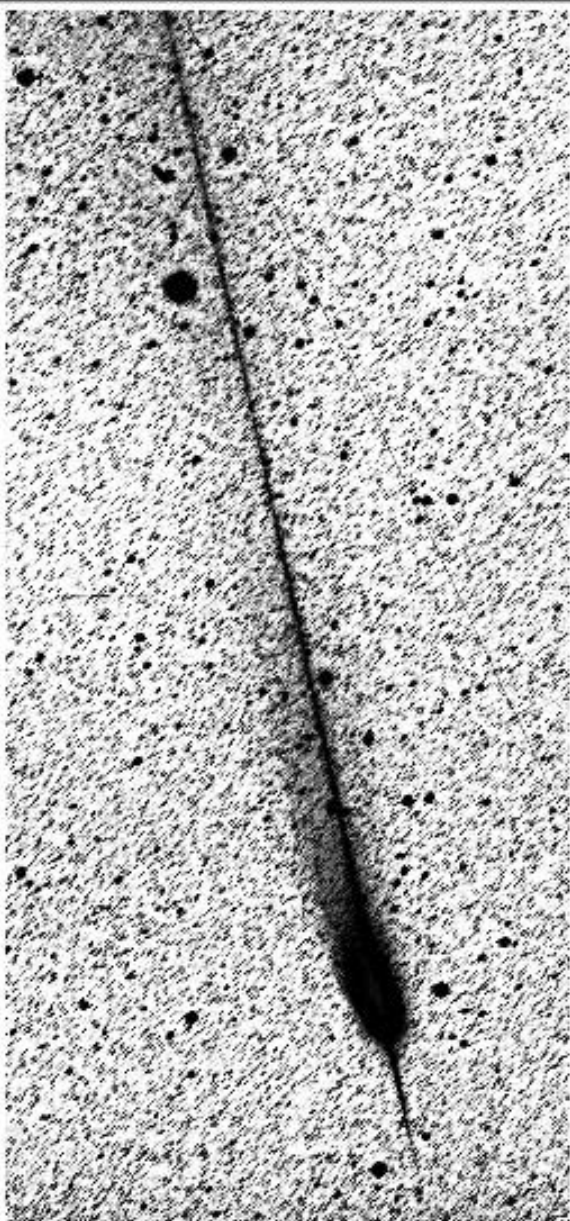
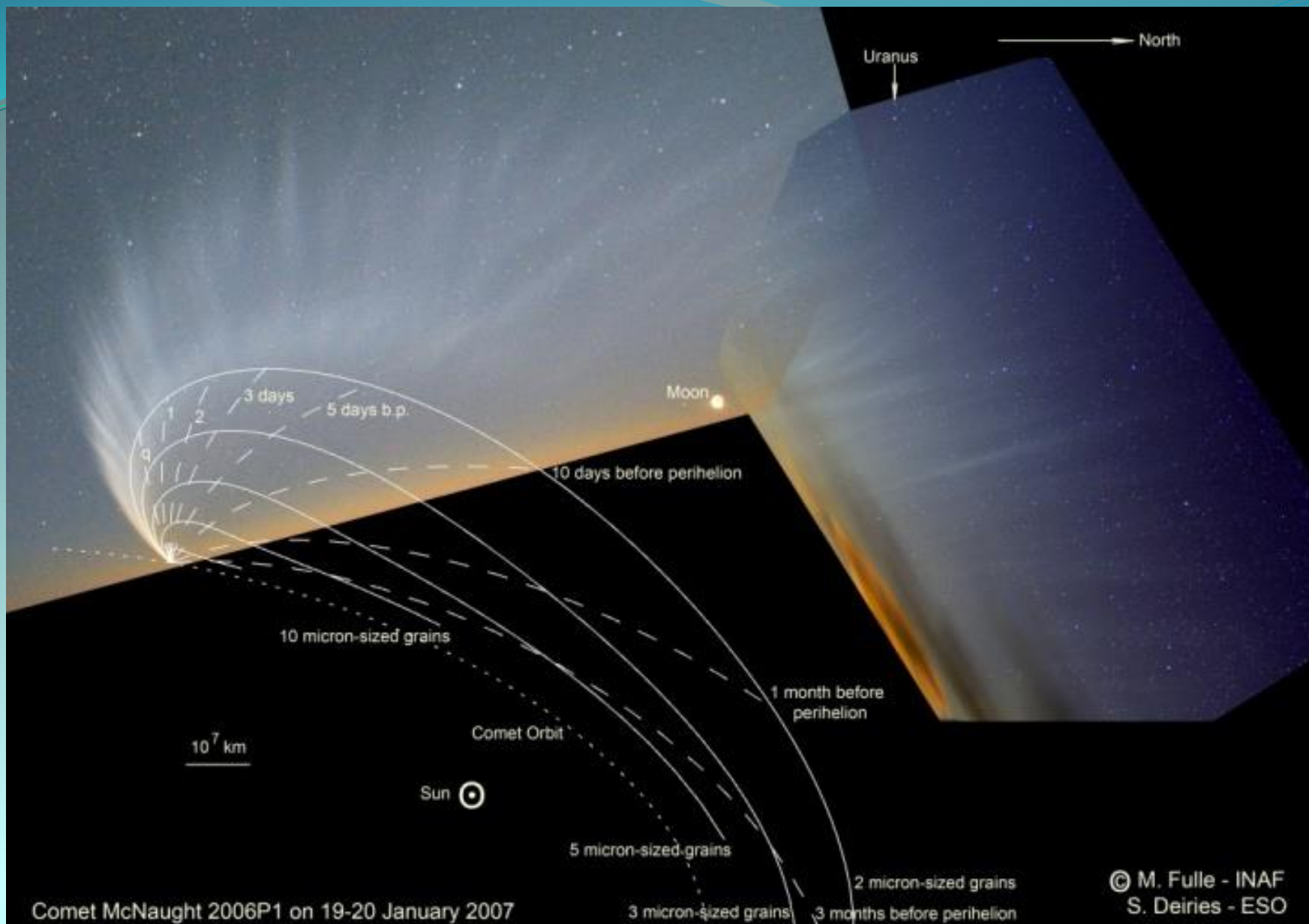


Fig. 2. Neck-line observed in Comet Hale-Bopp 1995O1 on 5 January 1998 with the ESO 1-m Schmidt Telescope. The original image was filtered (unsharp masking) to enhance the spike features: the real antitail pointing toward the Sun (bottom) and the neck-line pointing in the opposite direction. ESO Press Photo 05a/98, courtesy of the European Southern Observatory (observer Guido Pizarro).



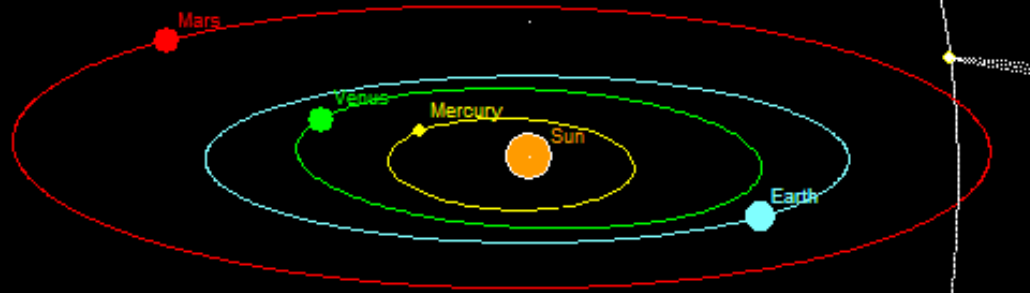
STRUTTURE DI POLVERI : LA “NECK LINE”





S. Deiries – M. Fulle

C/2018 W2 (AFRICANO)



Days to Perihelion : 2.896880
Now : 8/9/2019
Phase angle (°) : 41.172055
True anomaly (°) : 2.141521
Earth distance (AU) : 0.779465
Sun distance (AU) : 1.453768

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